#### Minimal Quantities and Measurability in Quantum Theory and Thermodynamics. Some Implications

Alexander Shalyt-Margolin<sup>1</sup>

Research Institute for Nuclear Problems, Belarusian State University, 11 Bobruiskaya str., Minsk 220040, Belarus

PACS: 03.65, 05.20 Keywords: minimal length, minimal inverse temperature, measurability

#### Abstract

In thermodynamics the measurability notion, introduced previously in a quantum theory, is defined on the basis of a minimal inverse temperature. Based on this notion, some implications are obtained for thermodynamics of black holes at all the energy scales and for quantum corrections of the basic quantities in the general case.

## 1 Introduction.

This paper is a continuation of the earlier works published by the author [1], [2] and [3]-[11]. The main target of these works is to construct a correct quantum theory and gravity in terms of the variations (increments) dependent on the existent energies.

At the present time physics is using (not without success) the mathematical apparatus based on the infinitesimal space-time variations (increments)

$$dt, dx_i, i = 1, \dots, 3 \tag{1}$$

This mathematical apparatus comes from mathematical analysis [12], [13] and is completely adequate for classical mechanics [14],[15], where continuous space-time forms the base.

<sup>&</sup>lt;sup>1</sup>E-mail: a.shalyt@mail.ru; alexm@hep.by

<sup>1</sup> 

But in this approach, due to the introduction of ultraviolet and infrared divergences into a Quantum Theory (QT) [16] and also due to the absence of correct passage to the high-energy (ultraviolet) region in Gravity (GR)[17], we are facing very serious problems.

By the author's opinion, these problems are solvable but beyond the paradigm of continuous space-time.

To solve this problems, in the above-mentioned works, using the minimal length  $l_{min}$  (minimal time  $t_{min}$ ), the author investigates a discrete space-time model, for which at low energies (far from the Planck energies) the results are to a high a curacy identical to those obtained with a continuous space-time model. And at high (Planck's) energies the indicated model is fundamentally discrete, leading to principally new results. All variations in any physical system considered in such a discrete model should be dependent on the existent energies.

The primary instrument for such a discrete model is the **measurability** notion introduced in [1], [2].

This paper demonstrates that a similar *(in essence dual)* notion may be also introduced in thermodynamics on the basis of a minimal inverse temperature, leading to very interesting inferences for thermodynamics of black holes at all the energy scales and for quantum corrections of the basic quantities.

Actually, all the required preliminary information is included in this text to gain a better understanding by the reader.

Section 2 presents in detail some of the results, obtained by the author in thermodynamics and published in [18], which are important in what follows. In the first part of Section 3 (Subsection 3.1.) the author gives consideration to the principal definitions from [1],[2] which are used for the derivation of the necessary formulae in the second part of this Section. In the second part of Section 3 (Subsection 3.2.) the **measurability** notion in thermodynamics is introduced; some direct inferences are drawn. Finally, in Subsection 4 the earlier obtained results are used to study thermodynamics of black holes at all the energy scales and to derive quantum corrections of the basic quantities.

# 2 Generalized Uncertainty Principles in Quantum Theory and Thermodynamics

In this Subsection the author presents some of the results from Section 2 of the paper [18], because they are important for this work.

It is well known that in thermodynamics an inequality for the pair interior energy - inverse temperature that is completely analogous to the standard uncertainty relation in quantum mechanics [19] can be written [22] – [27]. The only (but essential) difference of this inequality from the quantum mechanical one is that the main quadratic fluctuation is defined by means of the classical partition function rather than by the quantum mechanical expectation values. In the last years a lot of papers appeared in which the usual momentum-coordinate uncertainty relation has been modified at very high energies of order Planck energy  $E_p$  [28]–[39]. In this note we propose simple reasons for modifying the thermodynamic uncertainty relation at Planck energies. This modification results in existence of the minimal possible main quadratic fluctuation of the inverse temperature. Of course we assume that all the thermodynamic quantities used are properly defined so that they have physical sense at such high energies.

We start with usual Heisenberg Uncertainty Principle (relation) [19] for momentum - coordinate:

$$\Delta x \ge \frac{\hbar}{\Delta p}.\tag{2}$$

It was shown that at the Planck scale the high-energy term should be as follows:

$$\Delta x \ge \frac{\hbar}{\Delta p} + \alpha' l_p^2 \frac{\Delta p}{\hbar} \tag{3}$$

where  $l_p$  is the Planck length  $l_p^2 = G\hbar/c^3 \simeq 1, 6 \ 10^{-35}m$  and  $\alpha'$  is a constant. In [28] this term is derived from the string theory, in [31] it results from the simple estimates of the Newtonian gravity and quantum mechanics, in [35] it comes from the black hole physics, other methods can also be used [34],[36],[37]. Relation (3) is quadratic in  $\Delta p$ 

$$\alpha' l_p^2 \left(\Delta p\right)^2 - \hbar \,\Delta x \Delta p + \hbar^2 \le 0 \tag{4}$$

and therefore leads to the fundamental length

$$\Delta x_{min} = 2\sqrt{\alpha' l_p} \tag{5}$$

Inequality (3) is called the Generalized Uncertainty Principle (GUP) in Quantum Theory.

Using relations (3), we can easily obtain a similar relation for the energy time pair. Indeed, (3) gives

$$\frac{\Delta x}{c} \ge \frac{\hbar}{\Delta pc} + \alpha' l_p^2 \frac{\Delta p}{c\hbar},\tag{6}$$

then

$$\Delta t \ge \frac{\hbar}{\Delta E} + \alpha' \frac{l_p^2}{c^2} \frac{\Delta pc}{\hbar} = \frac{\hbar}{\Delta E} + \alpha' t_p^2 \frac{\Delta E}{\hbar}.$$
(7)

where the smallness of  $l_p$  is taken into account so that the difference between  $\Delta E$  and  $\Delta(pc)$  can be neglected and  $t_p$  is the Planck time  $t_p = l_p/c = \sqrt{G\hbar/c^5} \simeq 0.54 \ 10^{-43} sec$ . Inequality (7) gives, similar to (3), the lower boundary for time  $\Delta t \geq 2t_p$  determining the fundamental time

$$t_{min} = 2\sqrt{\alpha'} t_p \tag{8}$$

Thus, we can write the inequalities in the standard form

$$\begin{cases} \Delta x \geq \frac{\hbar}{\Delta p} + \alpha' \left(\frac{\Delta p}{P_{pl}}\right) \frac{\hbar}{P_{pl}} \\ \Delta t \geq \frac{\hbar}{\Delta E} + \alpha' \left(\frac{\Delta E}{E_p}\right) \frac{\hbar}{E_p} \end{cases}$$
(9)

where  $P_{pl} = E_p/c = \sqrt{\hbar c^3/G}$ . Now we consider the thermodynamic uncertainty relations between the inverse temperature and interior energy of a macroscopic ensemble

$$\Delta \frac{1}{T} \ge \frac{k_B}{\Delta U}.\tag{10}$$

where  $k_B$  is the Boltzmann constant.

N.Bohr [20] and W.Heisenberg [21] first pointed out that such a kind of uncertainty principle is applicable in thermodynamics. The thermodynamic

uncertainty relations (10) were proven by many authors in various ways [22] - [27]. Therefore, their validity is unquestionable. Nevertheless, relation (10) was proven in view of the standard model of the infinite-capacity heat bath encompassing the ensemble. But it is obvious from the above inequalities that at very high energies the capacity of the heat bath can no longer be assumed infinite at the Planck scale. Indeed, the total energy of the pair heat bath - ensemble may be arbitrary large but finite merely as the universe is born at a finite energy. Hence the quantity that can be interpreted as a temperature of the ensemble must have the upper limit and so does its main quadratic deviation. In other words, the quantity  $\Delta(1/T)$  must be bounded from below. But in this case an additional term should be introduced into (10)

$$\Delta \frac{1}{T} \ge \frac{k_B}{\Delta U} + \eta \,\Delta U \tag{11}$$

where  $\eta$  is a coefficient. Dimension and symmetry reasons give

$$\eta \sim \frac{k_B}{E_p^2} \quad or \quad \eta = \alpha' \frac{k_B}{E_p^2}$$

$$\tag{12}$$

As in the previous cases, inequality (11) leads to the fundamental (inverse) temperature.

$$T_{max} = \frac{\hbar}{2\sqrt{\alpha' t_p k_B}} = \frac{E_p}{2\sqrt{\alpha' k_B}} = \frac{T_p}{2\sqrt{\alpha'}} = \frac{\hbar}{t_{min} k_B},$$
$$\beta_{min} = \frac{1}{k_B T_{max}} = \frac{t_{min}}{\hbar}$$
(13)

In [40] the black hole horizon temperature is measured with the use of the Gedanken experiment. In the process the Generalized Uncertainty Relations in Thermodynamics (11) are derived too. Expression (11) is considered in the monograph [41] within the scope of the mathematical physics methods. Thus, we obtain a system of the generalized uncertainty relations in a sym-

metric form

$$\begin{aligned}
\Delta x &\geq \frac{\hbar}{\Delta p} + \alpha' \left(\frac{\Delta p}{P_{pl}}\right) \frac{\hbar}{P_{pl}} + \dots \\
\Delta t &\geq \frac{\hbar}{\Delta E} + \alpha' \left(\frac{\Delta E}{E_p}\right) \frac{\hbar}{E_p} + \dots \\
\Delta \frac{1}{T} &\geq \frac{k_B}{\Delta U} + \alpha' \left(\frac{\Delta U}{E_p}\right) \frac{k_B}{E_p} + \dots
\end{aligned}$$
(14)

or in the equivalent form

$$\begin{cases} \Delta x \geq \frac{\hbar}{\Delta p} + \alpha' l_p^2 \frac{\Delta p}{\hbar} + \dots \\ \Delta t \geq \frac{\hbar}{\Delta E} + \alpha' t_p^2 \frac{\Delta E}{\hbar} + \dots \\ \Delta \frac{1}{T} \geq \frac{k_B}{\Delta U} + \alpha' \frac{1}{T_p^2} \frac{\Delta U}{k_B} + \dots \end{cases}$$
(15)

where the dots mean the existence of higher-order corrections as in [42]. Here  $T_p$  is the Planck temperature:  $T_p = E_p/k_B$ . (4)

In literature the relation (10) is referred to as the Uncertainty Principle in Thermodynamics (UPT). Let us call relation (11) the Generalized Uncertainty Principle in Thermodynamics (GUPT).

In this case, without the loss of generality and for symmetry, it is assumed that a dimensionless constant in the right-hand side of GUP (formula (3)) and in the right-hand side of GUPT (formula (11)) is the same  $-\alpha'$ .

# 3 Minimal Length, Minimal Inverse Temperature, and Measurability in Quantum Theory and Thermodynamics

### 3.1 Minimal Length and Measurability Notion in Quantum Theory

First, we consider in this Subsection the principal definitions from [1],[2] which are required to derive the key formulae in the second part of the Subsection and to obtain further results.

**Definition I.** Let us call as **primarily measurable variation** any small variation (increment)  $\widetilde{\Delta}x_{\mu}$  of any spatial coordinate  $x_{\mu}$  of the arbitrary point  $x_{\mu}, \mu = 1, ..., 3$  in some space-time system R if it may be realized in the form of the uncertainty (standard deviation)  $\Delta x_{\mu}$  when this coordinate is measured within the scope of Heisenberg's Uncertainty Principle (HUP) [19] (formula (2) in general case):

$$\widetilde{\Delta}x_{\mu} = \Delta x_{\mu}, \Delta x_{\mu} \simeq \frac{\hbar}{\Delta p_{\mu}}, \mu = 1, 2, 3$$
(16)

for some  $\Delta p_{\mu} \neq 0$ .

Similarly, for  $\mu = 0$  for pair "time-energy" (t, E), let's call any small variation (increment) by **primarily measurable variation** in the value of time  $\widetilde{\Delta}x_0 = \widetilde{\Delta}t_0$  if it may be realized in the form of the uncertainty (standard deviation)  $\Delta x_0 = \Delta t$  and then

$$\widetilde{\Delta}t = \Delta t, \Delta t \simeq \frac{\hbar}{\Delta E} \tag{17}$$

for some  $\Delta E \neq 0$ . Formula (17) is nothing else as formula (7) for  $\Delta E \ll E_p$ Here HUP is given for the nonrelativistic case. In the relativistic case HUP has the distinctive features [43] which, however, are of no significance for the general formulation of **Definition I.**, being associated only with particular alterations in the right-hand side of the second relation Equation (17). It is clear that at low energies  $E \ll E_P$  (momenta  $P \ll P_{pl}$ ) **Definition I.** 

sets a lower bound for the **primarily measurable variation**  $\Delta x_{\mu}$  of any space-time coordinate  $x_{\mu}$ .

At high energies E (momenta P) this is not the case if E(P) have no upper limit. But, according to the modern knowledge, E(P) are bounded by some maximal quantities  $E_{max}$ ,  $(P_{max})$ 

$$E \le E_{max}, P \le P_{max},\tag{18}$$

where in general  $E_{max}$ ,  $P_{max}$  may be on the order of Planck quantities  $E_{max} \propto E_P$ ,  $P_{max} \propto P_{pl}$  and also may be the trans-Planck's quantities.

In any case the quantities  $P_{max}$  and  $E_{max}$  lead to the introduction of the minimal length  $l_{min}$  and of the minimal time  $t_{min}$ .

Supposition II. There is the minimal length  $l_{min}$  as a minimal measurement unit for all primarily measurable variations having the dimension of length, whereas the minimal time  $t_{min} = l_{min}/c$  as a minimal measurement unit for all quantities or primarily measurable variations (increments) having the dimension of time, where c is the speed of light.

 $l_{min}$  and  $t_{min}$  are naturally introduced as  $\Delta x_{\mu}$ ,  $\mu = 1, 2, 3$  and  $\Delta t$  in Equations (16) and (17) for  $\Delta p_{\mu} = P_{max}$  and  $\Delta E = E_{max}$ .

For definiteness, we consider that  $E_{max}$  and  $P_{max}$  are the quantities on the order of the Planck quantities, then  $l_{min}$  and  $t_{min}$  are also on the order of Planck quantities  $l_{min} \propto l_P$ ,  $t_{min} \propto t_P$ .

**Definition I.** and **Supposition II.** are quite natural in the sense that there are no physical principles with which they are inconsistent.

The combination of **Definition I.** and **Supposition II.** will be called the **Principle of Bounded Primarily Measurable Space-Time Varia**tions (Increments) or for short **Principle of Bounded Space-Time Variations (Increments)** with abbreviation (PBSTV).

As the minimal unit of measurement  $l_{min}$  is available for all the **primarily** measurable variations  $\Delta L$  having the dimensions of length, the "Integrality Condition" (IC) is the case

$$\Delta L = N_{\Delta L} l_{min},\tag{19}$$

where  $N_{\Delta L} > 0$  is an integer number.

In a like manner the same "Integrality Condition" (IC) is the case for all

the **primarily measurable variations**  $\Delta t$  having the dimensions of time. And similar to Equation (19), we get the for any time  $\Delta t$ :

$$\Delta t \equiv \Delta t(N_t) = N_{\Delta t} t_{min},\tag{20}$$

where similarly  $N_{\Delta t} > 0$  is an integer number too.

Definition 1 (Primary or Elementary Measurability.)

In accordance with the PBSTV let us define the quantity having the dimensions of length or time as primarily (or elementarily) measurable, when it satisfies the relation Equation (19) (and respectively Equation (20)).
 Let us define any physical quantity primarily (or elementarily) measurable, when its value is consistent with points (1) of this Definition.

It is convenient to use the deformation parameter  $\alpha_a$ . This parameter has been introduced earlier in the papers [44],[18],[45]–[48] as a *deformation parameter* (in terms of paper [49]) on going from the canonical quantum mechanics to the quantum mechanics at Planck's scales (Early Universe) that is considered to be the quantum mechanics with the minimal length (QMML):

$$\alpha_a = l_{min}^2 / a^2, \tag{21}$$

where a is the measuring scale. It is easily seen that the parameter  $\alpha_a$  from Equation (21) is discrete as it is nothing else but

$$\alpha_a = l_{min}^2 / a^2 = \frac{l_{min}^2}{N_a^2 l_{min}^2} = \frac{1}{N_a^2}.$$
(22)

At the same time, from Equation (22) it is evident that  $\alpha_a$  is irregularly discrete.

It should be noted that, physical quantities complying with **Definition 1** won't be enough for the research of physical systems. Indeed, such a variable as

$$\alpha_{N_a l_{min}}(N_a l_{min}) = p(N_a) \frac{l_{min}^2}{\hbar} = l_{min}/N_a, \qquad (23)$$

(where  $\alpha_{N_a l_{min}} = \alpha_a$  is taken from formula (22) at  $a = N_a l_{min}$ , and  $p(N_a) = \frac{\hbar}{N_a l_{min}}$  is the corresponding **primarily measurable** momentum), is fully expressed in terms *only* **Primarily Measurable Quantities** of **Definition** 

1 and that's why it may appear at any stage of calculations, but apparently doesn't comply with **Definition 1**. That's why it's necessary to introduce the following definition generalizing **Definition 1**:

#### Definition 2. Generalized Measurability

We shall call any physical quantity as **generalized-measurable** or for simplicity **measurable** if any of its values may be obtained in terms of **Primarily Measurable Quantities** of **Definition 1**.

In what follows, for simplicity, we will use the term **Measurability** instead of **Generalized Measurability**.

It is evident that any **primarily measurable quantity (PMQ)** is **measurable**. Generally speaking, the contrary is not correct, as indicated by formula (23).

The generalized-measurable quantities are appeared from the Generalized Uncertainty Principle (GUP) (formula (3)) that naturally leads to the minimal length  $l_{min}$  [28]–[39]:

$$\Delta x_{min} = 2\sqrt{\alpha' l_p} \doteq l_{min},\tag{24}$$

For convenience, we denote the minimal length  $l_{min} \neq 0$  by  $\ell$  and  $t_{min} \neq 0$  by  $\tau = \ell/c$ .

Solving inequality (3), in the case of equality we obtain the apparent formula

$$\Delta p_{\pm} = \frac{(\Delta x \pm \sqrt{(\Delta x)^2 - 4\alpha' l_p^2})\hbar}{2\alpha' l_p^2}.$$
(25)

Next, into this formula we substitute the right-hand part of formula (19) for L = x. Considering (24), we can derive the following:

$$\Delta p_{\pm} = \frac{(N_{\Delta x} \pm \sqrt{(N_{\Delta x})^2 - 1})\hbar\ell}{\frac{1}{2}\ell^2} = \frac{2(N_{\Delta x} \pm \sqrt{(N_{\Delta x})^2 - 1})\hbar}{\ell}.$$
(26)

But it is evident that at low energies  $E \ll E_p$ ;  $N_{\Delta x} \gg 1$  the plus sign in the nominator (26) leads to the contradiction as it results in very high (much

greater than the Planck's) values of  $\Delta p$ . Because of this, it is necessary to select the minus sign in the numerator (26). Then, multiplying the left and right sides of (26) by the same number  $N_{\Delta x} + \sqrt{N_{\Delta x}^2 - 1}$ , we get

$$\Delta p = \frac{2\hbar}{(N_{\Delta x} + \sqrt{N_{\Delta x}^2 - 1})\ell}.$$
(27)

 $\Delta p$  from formula (27) is the **generalized-measurable** quantity in the sense of **Definition 2.** However, it is clear that at low energies  $E \ll E_p$ , i.e. for  $N_{\Delta x} \gg 1$ , we have  $\sqrt{N_{\Delta x}^2 - 1} \approx N_{\Delta x}$ . Moreover, we have

$$\lim_{N_{\Delta x} \to \infty} \sqrt{N_{\Delta x}^2 - 1} = N_{\Delta x}.$$
 (28)

Therefore, in this case (27) may be written as follows:

$$\Delta p \doteq \Delta p(N_{\Delta x}, HUP) = \frac{\hbar}{1/2(N_{\Delta x} + \sqrt{N_{\Delta x}^2 - 1})\ell} \approx \frac{\hbar}{N_{\Delta x}\ell} = \frac{\hbar}{\Delta x};$$

$$N_{\Delta x} \gg 1, \qquad (29)$$

in complete conformity with HUP. Besides,  $\Delta p \doteq \Delta p(N_{\Delta x}, HUP)$ , to a high accuracy, is a **primarily measurable** quantity in the sense of **Definition 1**.

And vice versa it is obvious that at high energies  $E \approx E_p$ , i.e. for  $N_{\Delta x} \approx 1$ , there is no way to transform formula (27) and we can write

$$\Delta p \doteq \Delta p(N_{\Delta x}, GUP) = \frac{\hbar}{1/2(N_{\Delta x} + \sqrt{N_{\Delta x}^2 - 1})\ell}; N_{\Delta x} \approx 1.$$
(30)

At the same time,  $\Delta p \doteq \Delta p(N_{\Delta x}, GUP)$  is a **Generalized Measurable** quantity in the sense of **Definition 2**. Thus, we have

$$GUP \to HUP$$
 (31)

for

$$(N_{\Delta x} \approx 1) \to (N_{\Delta x} \gg 1).$$
 (32)

Also, we have

$$\Delta p(N_{\Delta x}, GUP) \to \Delta p(N_{\Delta x}, HUP), \tag{33}$$

where  $\Delta p(N_{\Delta x}, GUP)$  is taken from formula (30), whereas  $\Delta p(N_{\Delta x}, HUP)$  from formula (29).

Comment  $2^*$ .

From the above formulae it follows that, within GUP, the **primarily measurable** variations (quantities) are derived to a high accuracy from the **generalized-measurable** variations (quantities) only in the low-energy limit  $E \ll E_P$ 

Next, within the scope of GUP, we can correct a value of the parameter  $\alpha_a$  from formula (22) substituting *a* for  $\Delta x$  in the expression  $1/2(N_{\Delta x} + \sqrt{N_{\Delta x}^2 - 1})\ell$ .

Then at low energies  $E \ll E_p$  we have the **primarily measurable** quantity  $\alpha_a(HUP)$ 

$$\alpha_a \doteq \alpha_a(HUP) = \frac{1}{[1/2(N_a + \sqrt{N_a^2 - 1})]^2} \approx \\ \approx \frac{1}{N_a^2}; N_a \gg 1,$$
(34)

that corresponds, to a high accuracy, to the value from formula (22). Accordingly, at high energies we have  $E \approx E_p$ 

$$\alpha_a \doteq \alpha_a (GUP) = \frac{1}{[1/2(N_a + \sqrt{N_a^2 - 1})]^2}; N_a \approx 1.$$
(35)

When going from high energies  $E \approx E_p$  to low energies  $E \ll E_p$ , we can write

$$\alpha_a(GUP) \xrightarrow{(N_a \approx 1) \to (N_a \gg 1)} \alpha_a(HUP) \tag{36}$$

in complete conformity to Comment  $2^*$ .

### 3.2 Minimal Inverse Temperature and Measurability

Now, let us return to the thermodynamic relation (11) in the case of equality:

$$\Delta \frac{1}{T} = \frac{k_B}{\Delta U} + \eta \,\Delta U,\tag{37}$$

that is equivalent to the quadratic equation

$$\eta \left(\Delta U\right)^2 - \Delta \frac{1}{T} \Delta U + k_B = 0.$$
(38)

The discriminant of this equation, with due regard for formula (12), is equal to

$$D = (\Delta \frac{1}{T})^2 - 4\eta k_B = (\Delta \frac{1}{T})^2 - 4\alpha' \frac{k_B^2}{E_p^2} \ge 0,$$
(39)

leading directly to  $(\Delta \frac{1}{T})_{min}$ 

$$(\Delta \frac{1}{T})_{min} = 2\sqrt{\alpha' \frac{k_B}{E_p}} \tag{40}$$

or, due to the fact that  $k_B$  is constant, we have

$$(\Delta \frac{1}{k_B T})_{min} = \frac{2\sqrt{\alpha'}}{E_p}.$$
(41)

It is clear that  $(\Delta \frac{1}{T})_{min}$  corresponds to  $T_{max}$  from formula (13)

$$T_{max} \approx T_p \gg 0. \tag{42}$$

In this case  $\Delta \frac{1}{T} \approx \frac{1}{T}$  and, of course, we can assume that

$$(\frac{1}{T})_{min} \doteq \tilde{\tau} = \frac{1}{T_{max}}.$$
(43)

Trying to find from formula (43) a minimal unit of measurability for the inverse temperature and introducing the "Integrality Condition" (IC) in line with the conditions (19),(20)

$$\frac{1}{T} = N_{1/T} \widetilde{\tau},\tag{44}$$

where  $N_{1/T} > 0$  is an integer number, we can introduce an analog of the **primary measurability** notion into thermodynamics.

#### **Definition 3** (Primary Thermodynamic Measurability)

(1) Let us define a quantity having the dimensions of inverse temperature as primarily measurable when it satisfies the relation (44).
(2)Let us define any physical quantity in thermodynamics as primarily

measurable when its value is consistent with point (1) of this Definition.

**Definition 3** in thermodynamics is analogous to the **Primary Measurability** in a quantum theory (**Definition 1**).

Now we consider the quadratic equation (38) in terms of **measurable** quantities in the sense of **Definition 3**. In accordance with this definition and with formula (44)  $\Delta(1/T)$ , we can write

$$\Delta \frac{1}{T} = N_{\Delta(1/T)} \tilde{\tau}, \tag{45}$$

where  $N_{\Delta(1/T)} > 0$  is an integer number.

The quadratic equation (38) takes the following form:

$$\eta \, (\Delta U)^2 - N_{\Delta(1/T)} \widetilde{\tau} \Delta U + k_B = 0. \tag{46}$$

Then, due to formula (41), we can find the "**measurable**" roots of equation (46) for  $\Delta U$  as follows:

$$(\Delta U)_{meas,\pm} = \frac{[N_{\Delta(1/T)} \pm \sqrt{N_{\Delta(1/T)}^2 - 1}]\tilde{\tau}}{2\eta} = \frac{2k_B [N_{\Delta(1/T)} \pm \sqrt{N_{\Delta(1/T)}^2 - 1}]\tilde{\tau}}{\tilde{\tau}^2} = \frac{2k_B [N_{\Delta(1/T)} \pm \sqrt{N_{\Delta(1/T)}^2 - 1}]}{\tilde{\tau}}.$$
(47)

The last line in (47) is associated with the obvious relation  $2\eta = \frac{\tilde{\tau}^2}{2k_B}$ . In this way we derive a complete analog of the corresponding relation (26)

from a quantum theory by replacement

$$\Delta p_{\pm} \Rightarrow \Delta U_{meas,\pm}; N_{\Delta x} \Rightarrow N_{\Delta(1/T)}; \hbar \Rightarrow k_B.$$
(48)

As, for low temperatures and energies,  $T \ll T_{max} \propto T_p$ , we have  $1/T \gg 1/T_p$  and hence  $\Delta(1/T) \gg 1/T_p$  and  $N_{\Delta(1/T)} \gg 1$ .

Next, in analogy with Subsection 3.1, in formula (47) we can have only the minus-sign root, otherwise, at sufficiently high  $N_{\Delta(1/T)} \gg 1$  for  $(\Delta U)_{meas,+}$  we can get  $(\Delta U)_{meas,+} \gg E_p$ . But this is impossible for low temperatures (energies).

On the contrary, the minus sign in (47) is consistent with high and low energies.

So, taking the root value in (47) corresponding to this sign and multiplying the nominator and denominator in (47) by  $N_{\Delta(1/T)} + \sqrt{N_{\Delta(1/T)}^2 - 1}$ , we obtain

$$(\Delta U)_{meas} = \frac{2k_B}{(N_{\Delta(1/T)} + \sqrt{N_{\Delta(1/T)}^2 - 1})\tilde{\tau}}$$
(49)

to have a complete analog of the corresponding relation from (27) in a quantum theory by substitution according to formula (48).

Then it is clear that, in analogy with Subsection 3.1, for low energies and temperatures  $N_{\Delta(1/T)} \gg 1$  (49) may be rewritten as

$$(\Delta U)_{meas} \doteq (\Delta U)_{meas} (T \ll T_{max}) =$$

$$= \frac{2k_B}{(N_{\Delta(1/T)} + \sqrt{N_{\Delta(1/T)}^2 - 1})\widetilde{\tau}} \approx$$

$$\approx \frac{k_B}{N_{\Delta(1/T)}\widetilde{\tau}}, N_{\Delta(1/T)} \gg 1, \qquad (50)$$

i.e. the Uncertainty Principle in Thermodynamics (UPT, formula (10)) is involved. In this case, due to the last formula,  $\Delta U_{meas}$  represents a **primarily measurable** thermodynamic quantity in the sense of **Definition 3** to a high accuracy.

Of course, at high energies the last term in the formula (50) is lacking and, for  $T \approx T_{max}$ ;  $N_{\Delta(1/T)} \approx 1$ , we have:

$$(\Delta U)_{meas} \doteq (\Delta U)_{meas} (T \approx T_{max}) = \frac{k_B}{1/2(N_{\Delta(1/T)} + \sqrt{N_{\Delta(1/T)}^2 - 1})\tilde{\tau}},$$

$$N_{\Delta(1/T)} \approx 1.$$
(51)

From (51) it follows that at high temperatures (energies)  $(\Delta U)_{meas}$  could hardly be a **primarily measurable** thermodynamic quantity. Because of this, it is expedient to use a counterpart of **Definition 2**.

#### Definition 4. Generalized Measurability in Thermodynamics

Any physical quantity in thermodynamics may be referred to as **generalized-measurable** or, for simplicity, **measurable** if any of its values may be obtained in terms of the **Primary Thermodynamic Measurability** of **Definition 3**.

In this way  $(\Delta U)_{meas}$  from the formula (51) is a **measurable** quantity. Based on the preceding formulae, it is clear that we have the limiting transition

$$(\Delta U)_{meas}(T \approx T_{max}) \xrightarrow{(N_{\Delta(1/T)} \approx 1) \to (N_{\Delta(1/T)} \gg 1)}$$
$$\xrightarrow{(N_{\Delta(1/T)} \approx 1) \to (N_{\Delta(1/T)} \gg 1)} (\Delta U)_{meas}(T \ll T_{max} \propto T_p),$$
(52)

that is analogous to the corresponding formula (36) in a quantum theory. Therefore, in this case the analog of *Comment 2\**. in Subsection 3.1 is valid. *Comment 2\* Thermodynamics* 

From the above formulae it follows that, within GUPT (11), the **primar**ily measurable variations (quantities) are derived, to a high accuracy, from the generalized-measurable variations (quantities) only in the lowtemperature limit  $T \ll T_{max} \propto T_p$ .

To conclude this Section, it seems logical to make several important **re-marks**.

**R3.1** It is obvious that all the calculations associated with **measurability** of inverse temperature  $\frac{1}{T}$  are valid for  $\beta = \frac{1}{k_B T}$  as well. Specifically, introducing  $\beta_{min} \doteq \tilde{\beta} = \tilde{\tau}/k_B$ , we can rewrite all the corresponding formulae in the "**measurable**" variant replacing 1/T ( $\Delta(1/T)$ ) by  $\beta, \tilde{\tau}$  by  $\tilde{\beta}$  and retaining  $N_{1/T}$  ( $N_{\Delta(1/T)}$ ).

**R3.2.** Naturally, the problem of compatibility between the **measurability** definitions in quantum theory and in thermodynamics arises: is there any contradiction between **Definition 1** from Subsection 3.1 and **Definitions 3** from Subsection 3.2?

On the basis of formulae (13) from Section 2 and (43) from Subsection 3.2 we can state:

measurability in quantum theory and thermodynamic measurability are completely compatible and consistent as the minimal unit of inverse temperature  $\tilde{\tau}$  is nothing else but the minimal time  $t_{min} = \tau$  up to a constant factor. And hence  $N_{1/T}$ ,  $(N_{\Delta(1/T)})$  is nothing else but  $N_t$ ,  $(N_{\Delta t})$  in (20). Then it is clear that  $N_t = N_{a=tc}$ .

**R3.3** Finally, from the above formulae (50), (51) it follows that the **measurable** temperature T is varying as follows:

$$T = \frac{T_{max}}{N_{1/T}}, T \ll T_{max} \propto T_p, N_{1/T} \gg 1;$$
  
$$T = \frac{T_{max}}{1/2(N_{1/T} + \sqrt{N_{1/T}^2 - 1})}, T \approx T_{max} \propto T_p,$$
  
$$N_{1/T} \approx 1.$$
 (53)

In such a way **measurable** temperature is a **discrete quantity** but at low energies it is almost constantly varying – so, the theoretical calculations are very similar to those of the well-known continuous theory. Actually, discreteness manifests itself in the case of high energies only.

# 4 Some Implications

### 4.1 Planck Deformation of Basic Quantities

The results from the previous Section may be interpreted as follows: at high energies i.e. at Planck's scales  $E \approx E_P$ , all the basic quantities l, t, T, and so on written by **measurable** terms from Sections 2,3 are modified (deformed) according to (30),(51).

In particular, we get

$$\begin{cases}
l \stackrel{(E \ll E_p) \to (E \approx E_p)}{\longrightarrow} \frac{1}{2} (l + \sqrt{l^2 - \ell^2}) \\
t \stackrel{(E \ll E_p) \to (E \approx E_p)}{\longrightarrow} \frac{1}{2} (t + \sqrt{t^2 - \tau^2}) \\
\frac{1}{T} \stackrel{(E \ll E_p) \to (E \approx E_p)}{\longrightarrow} \frac{1}{2} (\frac{1}{T} + \sqrt{(\frac{1}{T})^2 - \tilde{\tau}^2})
\end{cases}$$
(54)

In a similar way, by the use of formulae (30),(51) and of their inferences, we can obtain the *high-energy*  $(E \approx E_p)$  "measurable" deformation for all the other physical quantities P, E, U, ... initially specified at low energies  $E \ll E_p$  in terms of measurable quantities.

Consequently, we can derive the "measurable" quantum corrections"  $\Delta_Q$  for l, t, 1/T and so on:

$$\begin{cases}
\Delta_Q l = \frac{1}{2}(l + \sqrt{l^2 - \ell^2}) - l = \frac{1}{2}(\sqrt{l^2 - \ell^2} - l) \\
\Delta_Q t = \frac{1}{2}(t + \sqrt{t^2 - \tau^2}) - t = \frac{1}{2}(\sqrt{t^2 - \tau^2} - \tau) \\
\Delta_Q \frac{1}{T} = \frac{1}{2}(\sqrt{(\frac{1}{T})^2 - \tilde{\tau}^2} - \frac{1}{T}) \\
\dots \dots \dots \dots
\end{cases}$$
(55)

### 4.2 Classical and Quantum Schwarzschild's Black Hole

Now let us show the applicability of the results from Section 3 to a quantum theory of black holes. Consider the case of Schwarzschild's black hole.

It seems logical to support the idea suggested in the Introduction to the recent overview presented by seven authors [50]: "Since for (asymptotically flat Schwarzschild) black holes the temperatures increase as their masses decrease, soon after Hawking's discovery, it became clear that a complete description of the evaporation process would ultimately require a consistent quantum theory of gravity. This is necessary as the semiclassical formulation of the emission process breaks down during the final stages of the evaporation as characterized by Planckian values of the temperature and spacetime curvature". Naturally, it is important to study the transition from low to high energies in the indicated case.

In [1], [2] it is shown that at low energies  $E \ll E_p$  a "**measurable**" theory is very close to the initial continuous theory. As a result, in the case under study we can use the basic formulae from a continuous theory, considering them valid to a high accuracy.

In particular, in the notation used for large Schwarzschild's black hole [64]  $(E \ll E_p; N_a \gg 1)$ , we have in terms of **measurable** quantities for  $N_a = N_{r_s}$ 

$$r_s = N_{r_s}\ell = \frac{2GM}{c^2}; M = \frac{N_{r_s}\ell c^2}{2G} = \frac{\alpha_{r_s}^{-1/2}(HUP)\ell c^2}{2G}.$$
 (56)

As its temperature is given by the formula

$$T_H = \frac{\hbar c^3}{8\pi G M k_B},\tag{57}$$

at once we get

$$T_H = \frac{\hbar c}{4\pi k_B r_s} = \frac{\hbar c}{4\pi k_B N_{r_s} \ell} = \frac{\hbar c \alpha_{r_s}^{1/2} (HUP)}{4\pi k_B \ell}.$$
(58)

Comparing this expression to the expression with high  $N_{1/T}$   $(N_{1/T} \gg 1)$  for temperature from the equation (53) we can find that at low energies  $E \ll E_p$ , due to comment **R3.2.**, the number  $N_{1/T}$  is actually coincident with the number  $N_a$ :

$$N_{1/T} = N_{r_s} = \alpha_{r_s}^{-1/2} (HUP).$$
(59)

In this formalism for a "quantum" Schwarzschild black hole (i.e. at high

energies  $E \approx E_p; N_{r_s} \approx 1$ ) by virtue of the same eq.(53) formula (58) is replaced by

$$T_H(Q) = \frac{\hbar c}{4\pi k_B \frac{1}{2} (N_{r_s} + \sqrt{N_{r_s}^2 - 1})\ell} = \frac{\hbar c \alpha_{r_s}^{1/2} (GUP)}{4\pi k_B \ell}, \tag{60}$$

where

$$\alpha_{r_s}^{-1/2}(GUP) = \frac{1}{2}(N_{r_s} + \sqrt{N_{r_s}^2 - 1})$$
(61)

Similarly, from formula (56) we can calculate for the "measurable" mass M of a large Schwarzschild black hole its "measurable" quantity M(Q) in the "quantum" case:

$$M(Q) = \frac{1/2(N_{r_s} + \sqrt{N_{r_s}^2 - 1})\ell c^2}{2G} = \frac{\alpha_{r_s}^{-1/2}(GUP)\ell c^2}{2G}.$$
 (62)

In the low-energy case  $(E \ll E_p; N_{r_s} \gg 1)$  the well-known semiclassical Bekenstein-Hawking formula [64]

$$S_{Schw} = \frac{4\pi r_s^2}{4l_p^2} = \frac{\pi r_s^2}{l_p^2}$$
(63)

for entropy of a *large Schwarzschild black hole*, considering formula (5), may be written in terms of **measurable** quantities as follows:

$$S_{Schw,meas} = \frac{4\pi (N_{r_s}\ell)^2}{\ell^2/\alpha'} = 4\pi \alpha' N_{r_s}^2.$$
 (64)

Thereafter, in the case of high energies  $(E \approx E_p; N_{r_s} \approx 1)$  the semiclassical **"measurable"** Bekenstein-Hawking entropy  $S_{Schw,meas}$  (formula (64))for a "quantum" Schwarzschild black hole is replaced by the quantum entropy  $S_{Schw,meas-q}$  and we have

$$S_{Schw,meas-q} = \frac{4\pi (\frac{1}{2}(N_{r_s} + \sqrt{N_{r_s}^2 - 1})\ell)^2}{\ell^2/\alpha'} = \pi \alpha' (N_{r_s} + \sqrt{N_{r_s}^2 - 1})^2.$$
(65)

As noted in [1], the parameter  $\alpha_a = \alpha_a(HUP)$ , within constant factors, is coincident with the Gaussian curvature  $K_a$  [65] corresponding to **primary measurable**  $a = N_a \ell$ :

$$\alpha_a = \frac{\ell^2}{a^2} = \ell^2 K_a. \tag{66}$$

Because of this, the transition from  $\alpha_a(HUP)$  to  $\alpha_a(GUP)$  may be considered as a basis for "quantum corrections" to the Gaussian curvature  $K_a$  in the high-energy region  $E \approx E_p$ :

$$\alpha_a(GUP) - \alpha_a(HUP) =$$

$$= \ell^2 \left[ \frac{1}{1/4(N_a + \sqrt{N_a^2 - 1})^2 \ell^2} - \frac{1}{N_a^2 \ell^2} \right] =$$

$$= \ell^2 (K_a^Q - K_a), \quad (67)$$

where the "measurable quantum Gaussian curvature "  $K_a^Q$  is defined as

$$K_a^Q \doteq \frac{1}{1/4(N_a + \sqrt{N_a^2 - 1})^2 \ell^2}.$$
(68)

In a similar way we can derive a "measurable quantum correction" for the temperature T, for the mass M and entropy S of a Schwarzschild black hole:

$$\begin{cases} \Delta_Q T \doteq T_H(Q) - T_H = \frac{\hbar c}{4\pi k_B \ell} (\alpha_{r_s}^{1/2}(GUP) - \alpha_{r_s}^{1/2}(HUP)), \\ \Delta_Q M \doteq M(Q) - M = \frac{\ell c^2}{2G} (\alpha_{r_s}^{-1/2}(GUP) - \alpha_{r_s}^{-1/2}(HUP)) = \\ = \frac{\ell c^2}{4G} (\sqrt{N_{r_s}^2 - 1} - N_{r_s}) = \frac{c^2}{4G} (\sqrt{r_s^2 - \ell^2} - r_s), \\ \Delta_Q S \doteq S_{Schw,meas-q} - S_{Schw,meas} = \\ = \pi \alpha' (N_{r_s} + \sqrt{N_{r_s}^2 - 1})^2 - 4\pi \alpha' N_{r_s}^2 = \\ = \pi \alpha' (2N_{r_s} \sqrt{N_{r_s}^2 - 1} - 2N_{r_s}^2 - 1). \end{cases}$$

As indicated by the last formulae and by (55), the **measurable quantum** corrections are nothing else but the difference between the generalized measurable quantities and the primarily measurable quantities of Subsection 3.1.

#### Remark 4.2.1

It is readily seen that a minimal value of  $N_a = 1$  is unattainable because in formula (30) we can obtain a value of the length l that is below the minimum  $l < \ell$  for the momenta and energies above the maximal ones, and that is impossible. Thus, we always have  $N_a \ge 2$ . This fact was indicated in [44],[18], however, based on the other approach.

#### **Conflict of Interests**

The author declares that there is no conflict of interests regarding the publication of this paper.

#### Acknowledgment

This work was supported by the Belarusian Republication Foundation for Fundamental Research (project N(16-121)).

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